

1 Characterization of oscillatory motions in the stable 2 atmosphere of a deep valley

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4
5 revised version, submitted 21 September 2012

6 **Abstract** In a valley sheltered from strong synoptic effects, the dynamics of the valley
7 atmosphere at night is dominated by katabatic winds. In a stably stratified atmosphere, these
8 winds undergo temporal oscillations, whose frequency is given by $N \sin \alpha$ for an infinitely
9 long slope of constant slope angle α , N being the buoyancy frequency. Such an unsteady
10 flow in a stably stratified atmosphere may also generate internal gravity waves (IGWs).
11 The numerical study by Chemel et al. (Meteorol Atmos Phys 203:187-194, 2009) showed
12 that, in the stable atmosphere of a deep valley, the oscillatory motions associated with the
13 IGWs generated by katabatic winds are distinct from those of the katabatic winds. The IGW
14 frequency was found to be independent of α and about $0.8 N$. Their study did not consider
15 the effects of the background stratification and valley geometry on these results. The present
16 work extends this study by investigating those effects for a wide range of stratifications and
17 slope angles, through numerical simulations for a deep valley. The two oscillatory systems
18 are reproduced in the simulations. The frequency of the oscillations of the katabatic winds
19 is found to be equal to N times the sine of the maximum slope angle. Remarkably, the IGW
20 frequency is found to also vary as $C_w N$, with C_w in the range $0.7 - 0.95$. These values for
21 C_w are similar to those reported for IGWs radiated by any turbulent field with no dominant
22 frequency component. Results suggest that the IGW wavelength is controlled by the valley
23 depth.

24 **Keywords** Complex terrain · Internal gravity waves · Katabatic winds · Numerical
25 simulations · Stably stratified atmosphere

26 1 Introduction

27 Under weak synoptic forcing, the flow in a deep valley is driven by thermal circulations due
28 to the heating or cooling of the ground surface. The cooling of the ground surface produces

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a shallow layer of cool, stable air above it, inducing downslope flows also referred to as katabatic flows (e.g. [Simpson 1994](#); [Poulos and Zhong 2008](#)). As the cold air flows down the slopes, it accumulates over the centre of the valley, thereby leading to a pool or lake of stably stratified air ([Mori and Kobayashi 1996](#); [Whiteman et al. 2008](#)).

The characteristics of katabatic winds were first reported from field measurements during summer nights. The winds have a jet-like velocity profile, with maximum of the order of 5 m s^{-1} reached at about 10 m above the ground, depending upon stratification and surface stress, and a very stable thermal gradient, up to 0.1 K m^{-1} (e.g. [Whiteman 2000](#)). Observational studies reported katabatic flows of different nature, from highly turbulent to quasi-steady flows, depending on the Richardson number (see for instance [Gryning et al. 1985](#); [Helmis and Papadopoulos 1996](#); [Monti et al. 2002](#); [Bastin and Drobinski 2005](#); [Princevac et al. 2008](#); [Viana et al. 2010](#); [Mahrt et al. 2010](#)). A refined classification based on dimensional analysis and momentum balance was proposed by ([Mahrt 1982](#)) for stationary flow over a constant slope with simple friction. Whatever their nature, katabatic flows have been shown to undergo temporal oscillations, in both observational and numerical studies (e.g. [van Gorsel et al. 2004](#); [Fedorovich and Shapiro 2009](#)). The existence of these oscillations was first accounted for by [Fleagle \(1950\)](#) for an isothermal (and thus stably stratified) atmosphere. As the air flows down the slopes, it undergoes locally a cycle of compressional warming, deceleration (because of the isothermal atmosphere), cooling by the ground surface and further acceleration. [McNider \(1982\)](#) showed, using a simple model coupling the along-slope velocity component and the potential temperature of a fluid particle, that buoyancy effects are responsible for the oscillatory behaviour of the katabatic flows. For an infinitely long slope with a slope angle α (with respect to the horizontal) of constant value and a constant vertical gradient of potential temperature, the model predicts that the frequency of these oscillations is $N \sin \alpha$, where N is the buoyancy (or Brunt-Väisälä) frequency (the square of which is proportional to the vertical gradient of potential temperature).

It is well-known that a body oscillating at frequency Ω in a stably-stratified fluid with constant N generates internal gravity waves (IGWs) if $\Omega < N$ (e.g. [Lighthill 1978](#)). The frequency of these waves (equal to Ω) is denoted ω_w , hereafter for clarity. These waves have a very peculiar dispersion relation ([von Görtler 1943](#); [Mowbray and Rarity 1967](#)): their frequency does not depend on the modulus of the wave vector \mathbf{k} but only on the angle ϕ that this wave vector makes with respect to the horizontal. In the absence of rotation, the dispersion relation for IGWs is

$$\omega_w^2 = N^2 \cos^2 \phi. \quad (1)$$

When IGWs are generated by a turbulent field with no dominant frequency component, observations of the radiated IGW field reveal that the IGWs propagate at a fixed angle with respect to the horizontal, of about 45° . This implies, from the dispersion relation (1), that a very narrow range of frequencies, centred about $0.7 N$, is actually excited (e.g. [Wu 1969](#); [Cerasoli 1978](#); [Dohan and Sutherland 2003](#); [Taylor and Sarkar 2007](#)).

Few studies have dealt with the generation of IGWs by katabatic flows ([Mori and Kobayashi 1996](#); [Renfrew 2004](#); [Yu and Cai 2006](#); [Princevac et al. 2008](#); [Whiteman et al. 2008](#); [Viana et al. 2010](#)). The numerical study by [Chemel et al. \(2009\)](#) showed that, in the stable atmosphere of an idealized deep valley, the oscillatory motions associated with the IGWs generated by katabatic winds are distinct from those of the katabatic winds. [Chemel et al. \(2009\)](#) also found that the power spectrum of the IGWs is peaked for a ratio $\omega_w/N \approx 0.8$, close to that observed in stably stratified turbulence. Only one numerical simulation was considered in this study, and so the generality of the results was not assessed. The aim of the present work is to extend the study of [Chemel et al. \(2009\)](#) by investigating

the effects of the background stratification of the atmosphere and valley geometry on the characteristics of the IGW field.

For this purpose, we analyse a set of numerical simulations, performed with the Advanced Regional Prediction System (ARPS), for a wide range of stratifications and slope angles. The design of the simulations is described in Sect. 2. The general features of the katabatic and valley winds are briefly reported in Sect. 3 while a detailed analysis of the oscillatory motions is presented in Sect. 4. The influence of the initial ground surface temperature, background stratification of the atmosphere and valley geometry on the characteristics of the IGW field are discussed in Sect. 5. Conclusions are given in Sect. 6.

2 Design of the numerical simulations

2.1 The numerical model

The numerical simulations are performed with the ARPS numerical model (Xue et al. 2000). The ARPS model is a non-hydrostatic atmospheric model that is appropriate for scales ranging from a few metres to hundreds of kilometres. The model solves the compressible Navier-Stokes equations, which describe the dynamics of the flow, using a terrain-following coordinate system. It involves surface layer physics and a soil model. In the present study, the air is considered as dry, even though microphysical processes are also included in the ARPS model. Spatial derivatives are discretized with a centered fourth-order finite difference scheme on a staggered grid of Arakawa C type. Time integration is performed with a centered leapfrog time difference scheme using a mode-splitting time integration technique to deal with the acoustic modes. The turbulent kinetic energy (TKE) 1.5-order closure scheme (Deardorff 1980) is used to model the subgrid scales.

2.2 The topography of the valley

The valley is oriented south-north and is connected to a plain to the south so that an along-valley wind can develop. The analytical expression for the topography of the valley is given by (see for instance Rampanelli et al. 2004)

$$h(x, y) = H h_x(x) h_y(y), \quad (2)$$

where

$$h_x(x) = \begin{cases} 1, & |x| > S_x + V_x \\ \frac{1}{2} - \frac{1}{2} \cos\left(\pi \frac{|x| - V_x}{S_x}\right), & V_x \leq |x| \leq S_x + V_x, \\ 0, & |x| < V_x \end{cases} \quad (3)$$

and

$$h_y(y) = \frac{1}{2} + \frac{1}{2} \tanh\left(\frac{y - y_0}{S_y}\right). \quad (4)$$

H is the valley depth, S_x and S_y are the width of the sloping sidewall along the west-east direction x and south-north direction y , respectively, $2V_x$ is the width of the valley floor. Note that $h(x, y) = 0$ for $-V_x < x < V_x$, whatever y , implying that the valley floor is flat. The function $h_y(y)$, which defines the height of the plateaux along the valley axis, is displayed in Fig. 5.

113 We used two sets of values for these parameters, which correspond to the topographies
 114 referred to as T1 and T2 hereafter. The topography T1, used in Chemel et al. (2009), is
 115 characterized by a valley length of 20 km, with $H = 1700$ m, $S_x = 2640$ m, $S_y = 5000$ m,
 116 $V_x = 620$ m (and $y_0 = 10$ km since y varies between 0 and 20 km in the present paper)
 117 (see Fig. 1a). Defined by this set of parameters, the topography T1 can be considered as
 118 an idealized representation of the Chamonix valley, located in the French Alps. It is worth
 119 noting that, since the valley depth varies along the valley axis (see Fig. 1a), so does the
 120 maximum value of the valley side slope for a given x location. This maximum slope can be
 121 calculated from Equations 2 to 4, yielding $0.5 \pi h_y(y) / S_x$, corresponding to a slope angle of
 122 about 45° at the valley end. In the following, the words *the slope of the topography* or, more
 123 simply, *the slope* refers to the slope of the valley side wall.

124 The topography T2 is used in Sect. 5 to investigate the dependency of the IGW field on
 125 the geometry of the valley. The key difference between the topographies T1 and T2 is that
 126 the maximum slope angle for the topography T2 is approximately constant, at a value of
 127 about 30° , along the valley axis over a distance greater than half of the valley length. The
 128 valley length and the parameters S_x and V_x are the same as for the topography T1 but the
 129 sloping sidewall width along y and the valley depth are set to $S_y = 1200$ m and $H = 1000$ m,
 130 respectively.

131 2.3 Model setup

132 The model is run for a 3-hour nocturnal situation starting at 2200 UTC (corresponding to
 133 time $t = 0$) in winter at the latitude of the Chamonix valley. No katabatic flow is prescribed
 134 at the initial time. For a deep valley under stable conditions, as is the case here, the valley
 135 atmosphere is often decoupled from the air above the valley (see for instance Whiteman
 136 2000), and so no synoptic forcing was prescribed as well. The velocity field is thus set to
 137 zero in the numerical domain at the initial time.

138 2.3.1 The initial stratification

139 The initial buoyancy frequency N is set to a constant value, and so the initial vertical gradi-
 140 ent of potential temperature $d\theta/dz = (\theta_0/g) N^2$ is constant (i.e., the potential temperature
 141 increases linearly with height). The value of the reference potential temperature θ_0 is set to
 142 that of the initial near-surface potential temperature at the valley floor, namely 271 K. (The
 143 near-surface temperature is the temperature of the first grid point in the atmosphere above
 144 the ground.) Note that, in the study by Chemel et al. (2009), the initial buoyancy frequency
 145 profile was derived from measurements in the Riviera valley, located in the Swiss Alps, and
 146 varied with height. The constant value of N used in the present work will allow for a sensi-
 147 tivity study of the influence of the background stratification upon the IGW dynamics, by
 148 varying the initial value of N from 0.91×10^{-2} to 2.33×10^{-2} rad s $^{-1}$, corresponding to
 149 an initial stratification $d\theta/dz$ ranging from 2.3 to 15 K km $^{-1}$ (see Table 1). This range of
 150 values covers most stable situations encountered in a valley atmosphere.

151 2.3.2 The initial ground surface temperature and subsequent evolution

152 The temperature of the ground surface T_s (namely, the skin-surface temperature) is initial-
 153 ized with an offset from the temperature of the near-surface air T_a . The offset $T_s - T_a$ is set
 154 to either zero or -3 K depending on the simulation (see Table 1). The deep soil temperature

155 T_2 is initialized in a similar way with an offset $T_2 - T_a$ of value either zero or -5 K depending
 156 on the simulation (see Table 1). Note these conditions are imposed at the initial time only
 157 and are therefore not a continuous forcing.

158 The change of the temperature of the two soil layers with time is governed by a surface
 159 energy budget taking into account the radiational cooling of the surface from the emission of
 160 longwave radiation. The time evolution of the ground surface temperature T_s half way down
 161 the slope at $y = 15$ km for simulation S1 is displayed in Fig. 2. T_s decreases by a few K per
 162 hour (about 6 K h^{-1} during the first hour of the simulation and $1 - 2 \text{ K h}^{-1}$ afterwards).
 163 This rate of cooling is consistent with that derived from *in situ* measurements at a field site
 164 in Vermont, USA, reported by Peck (1996).

165 2.3.3 Boundary conditions

166 Open boundary conditions are used in the horizontal directions. **An impermeability condi-**
 167 **tion is imposed at the ground surface, namely, the velocity component normal to the ground**
 168 **is zero there.** A Rayleigh sponge is introduced at the top of the domain in order to absorb
 169 upward propagating waves.

170 The surface roughness length is set to 0.1 m, a value typical of cultivated areas. The
 171 Monin-Obukhov surface layer scheme is coupled to the two-layer soil-vegetation model
 172 developed by Noilhan and Planton (1989) to provide surface forcing in terms of momentum,
 173 heat and moisture fluxes.

174 2.4 Numerical parameters

175 The domain is discretized using 61×103 grid points in the horizontal, with a horizontal grid
 176 resolution of 200 m. The calculations are made on 140 vertical levels up to 7000 m. The grid
 177 mesh is stretched along the vertical to accommodate a high vertical resolution close to the
 178 ground surface, of 5 m below 100 m and then gradually increasing with height to reach 98 m
 179 at the top of the domain. The time step is 0.25 s.

180 3 The katabatic and valley winds for simulation S1

181 In this section and the next one, we focus on simulation S1 (see Table 1) for which the initial
 182 stratification is in the middle of the range of stratifications considered in our work.

183 3.1 The katabatic wind

184 3.1.1 General features of the katabatic wind

185 As is customary, we introduce a rotated coordinate system (s, n) where s is the coordinate
 186 along the sloping surface, positive down the slope, and n is the coordinate normal to the
 187 sloping surface, positive upwards. Note that the grid size along the n -axis at a given location
 188 along the slope is equal to $dn = \cos(\alpha_{loc}) dz$, where α_{loc} is the angle of the slope at this
 189 location. Since $dz = 5$ m in the present study, for $\alpha_{loc} = 45^\circ$ for instance, $dn \simeq 3.5$ m.

190 The velocity component along the sloping surface, denoted by u_s , is displayed in Fig. 3
 191 as a function of n for $y = 7$ km and $t = 74$ min, at the location of maximum slope angle,

192 equal here to 21° (i.e., half way down the slope). The katabatic wind is distributed in a layer
 193 immediately above the slope, of depth of about 30 m, and reaches a maximum value of
 194 $2 - 3 \text{ m s}^{-1}$ at the first grid point above the ground surface (i.e. at 2.5 m in our calculation).
 195 These features are in agreement with *in situ* measurements of katabatic winds on steep
 196 slopes (i.e., with a slope angle larger than 10°), either on a single slope (e.g. Helmis and
 197 Papadopoulos 1996; Monti et al. 2002) or in a valley (Gryning et al. 1985; van Gorsel
 198 et al. 2004). A return flow (of very small amplitude) is created above the downslope flow
 199 as a result of mass conservation. Slant observations of such a return flow are available (e.g.
 200 Buettner and Thyer 1965) as these require the katabatic wind to flow in a quiet environment
 201 and remain quasi-two-dimensional (that is, with no cross-slope wind). Such a return flow has
 202 also been observed in numerical simulations of katabatic flow (Skylingstad 2003; Catalano
 203 and Cenedese 2010).

204 3.1.2 Temporal oscillations of the katabatic wind

205 At a given location along the slope, the along-slope component of the wind u_s varies with
 206 time. This is attested in Fig. 4a where u_s is plotted near the bottom of the slope: u_s un-
 207 dergoes oscillations about a positive value of approximately 0.5 m s^{-1} , with an apparently
 208 well-defined period. These oscillations are present all along the slope. For an infinitely long
 209 slope of constant angle α and for a constant buoyancy frequency N , the frequency of these
 210 oscillations is given by (McNider 1982)

$$211 \quad \omega_k = N \sin \alpha. \quad (5)$$

212 For a given y location along the valley axis, α varies along the slope, and so it is not
 213 obvious which value should be used for α in Equation 5. The frequency spectrum associated
 214 with the time series of u_s displayed in Fig. 4a shows a dominant peak with a period of 10 min
 215 (see Fig. 4b). Using the value of N for simulation S1, Equation 5 yields a value for α of
 216 about 45° , which corresponds to the maximum angle of the slope at this y location. Hence,
 217 the period of the oscillations would be set by the background stratification and maximum
 218 angle of the slope. We show below that this result also holds for a more gentle slope.

219 In parts of the valley where the slope angle is smaller, the period of these oscillations
 220 becomes longer, in agreement with Equation 5. The frequency spectrum of u_s at the same x
 221 location as Fig. 4a but for $y = 7 \text{ km}$ (i.e., at a location closer to the valley mouth), shows
 222 a dominant peak for a period of 20 min (see Fig. 4c). For this period, Equation 5 leads to
 223 $\alpha = 21^\circ$, which is the maximum slope angle at that y location.

224 Note that the frequency spectra displayed in Fig. 4b and Fig. 4c show several peaks, the
 225 magnitudes of which vary with distance down the slope. It is still remarkable that the fluid
 226 particle model developed by McNider (1982) for an infinitely long slope with constant slope
 227 angle and a constant buoyancy frequency, predicts the dominant peak of these frequency
 228 spectra.

229 The oscillations of the katabatic wind have been mainly detected in a shallow layer
 230 immediately above the slopes, of depth 20 to 30 m. Having said that, oscillations have also
 231 been detected in the return flow and have a frequency close to that of the downslope flow
 232 (not shown). This indicates that the return flow is tightly coupled to the downslope flow.

233 3.2 The valley wind

234 As a result of downslope flows filling the valley with cold air, the valley atmosphere cools
 235 faster than the plain atmosphere at the same altitude, producing higher pressure in the valley
 236 compared to the plain. This pressure gradient drives a down-valley flow from the valley to
 237 the plain (see Fig. 5). As the cold air flowing down the slopes accumulates over the centre
 238 of the valley, the valley temperature inversion grows deeper and deeper. As the inversion
 239 layer deepens, the layer of down-valley wind deepens. As can be seen in Fig. 5, the depth of
 240 the inversion layer reaches that of the valley 45 min into the simulation. After the period of
 241 rapid growth of the valley inversion, the down-valley wind prevails through the depth of the
 242 fully developed inversion. It displays a minimum value of -0.5 m s^{-1} close to the ground, in
 243 a 40 m deep layer occurring close to the mouth of the valley.

244 The above description of the wind and temperature structure evolution in the valley
 245 agrees well with observations and can be found in textbooks (e.g. Whiteman 2000). How-
 246 ever, in our work, the establishment of the down-valley wind is a bit more subtle. Because
 247 the height of the plateaus and slope angle of the valley sidewalls increase with distance from
 248 the plain, the volume of air pouring down the slopes is larger as one moves toward the valley
 249 end. This results in an along-valley pressure gradient, which accelerates the flow.

250 4 The internal gravity wave field for simulation S1

251 4.1 Emission of the internal gravity wave field

252 Since the atmosphere is stably stratified, any non-horizontal perturbation varying in time
 253 with a frequency component smaller than N , generates IGWs. As shown in Sect. 3, katabatic
 254 winds are unsteady, with a frequency spectrum containing frequency components smaller
 255 than N (e.g. ω_k), and so should emit an IGW field propagating away from the slopes.

256 Let us show that Coriolis effects do not affect the wave dynamics. Once emitted, the
 257 wave dynamics can be assumed to be linear and, therefore, satisfy the dispersion relation.
 258 Accounting for Coriolis effects, relation (1) becomes $\omega_w^2 = N^2 \cos^2 \phi + f^2 \sin^2 \phi$, which can
 259 also be written as $(\omega_w/N)^2 = \cos^2 \phi + (f/N)^2 \sin^2 \phi$. For ϕ smaller than $\pi/2$, the second
 260 term of this dispersion relation can be neglected if $f/N \ll 1$. In the present case $f \simeq 10^{-4}$
 261 rad s^{-1} and $N = 1.47 \times 10^{-2} \text{ rad s}^{-1}$ (so that $f/N \simeq 0.007$) implying that rotation effects
 262 can be ignored. As discussed in classical textbooks (e.g. Lighthill 1978), in the absence
 263 of rotation, the flow induced by plane IGWs is a parallel shear flow, where the velocity is
 264 normal to the wave vector and lies in the same vertical plane. Hence the angle of the velocity
 265 vector with respect to the vertical is the angle ϕ in the dispersion relation (1).

266 The emission of IGWs by the unsteady katabatic winds is illustrated in Fig. 6b, in which
 267 the vertical velocity component w is displayed in a vertical cross section for $y = 15 \text{ km}$ and
 268 $t = 45 \text{ min}$. Since the generation of IGWs has just started (i.e., the wave-induced velocity is
 269 zero away from the slopes), the signature of the IGW field appears as upward and downward
 270 motions and resembles closed cells. The same feature was found in the numerical study by
 271 Renfrew (2004) for the IGW field generated by an unsteady (decelerating) katabatic flow on
 272 a slope-varying ice shelf (see Fig. 14 of this paper), and by Catalano and Cenedese (2010)
 273 when analysing nocturnal conditions in a valley of constant slope. A remarkable feature of
 274 this cell pattern is that the angle that the cell axis makes with the vertical is nearly constant
 275 along the valley sidewalls, despite the varying slope angle. The cell axis angle is ϕ which
 276 implies that the IGW frequency is constant for N constant (see the dispersion relation (1))

and independent of the slope angle. This important finding is further discussed in the next sections. The maximum value of the wave-induced vertical velocity is about 0.2 m s^{-1} , that is, one order of magnitude smaller than the vertical velocity of the katabatic flows, which emits the waves. (Indeed, using the relation $w = (\sin \alpha) u_s$ between the vertical and along-slope components of the velocity, one finds, for $\alpha = 45^\circ$ and from Fig. 3, that the maximum amplitude of w is about 1.5 m s^{-1} .)

Since the IGWs are generated by the katabatic flow, this flow needs to become established before the IGWs become apparent. For the conditions of simulation S1, about 20 min are required before IGWs can be observed. The wave field first appears at the bottom of the slopes (see Fig. 6a), as also found by Renfrew (2004) and Yu and Cai (2006). The latter authors, who conducted a numerical study similar to that of Renfrew (2004), focused on the most likely origin of the IGW field, namely the vertical velocity perturbation induced by a hydraulic jump in the katabatic flow. This hydraulic jump can result from the katabatic flow encountering a cold pool at the bottom of the valley or, simply, strongly decelerating as it reaches the foot of the slope. The possible occurrence of a hydraulic jump is usually estimated by computing a Froude number associated with the katabatic flow. Using the classical definition (e.g. Ball 1956) $Fr = U / [(g\Delta\theta/\theta_0)H]$, where U is a typical velocity of the katabatic flow and $\Delta\theta$ the potential temperature deficit across the katabatic wind layer of height H , one finds $Fr \simeq 2 - 3$ for simulation S1. Therefore $Fr > 1$ for the katabatic flow (and $Fr < 1$ downstream of that flow) so that a hydraulic jump is likely to occur. In the present case, this hydraulic jump would be created by the katabatic flow encountering the flat valley floor.

Fig. 6 shows that a standing IGW pattern is created at the bottom of the valley because of the convergence of the katabatic flows originating from the slopes of each valley sidewall. Finally, we note that the IGWs first form along the longest and steepest slopes (i.e., as one moves toward the valley end) and, later in time, along shorter and shallower slopes as well (see also Section 4.3 and Fig. 8a).

4.2 Frequency analysis of the internal gravity wave field

The purpose of this section is to show that a quasi-monochromatic IGW field develops, the frequency of which ω_w is independent of the frequency ω_k of the oscillations of the katabatic flows. In the following, we therefore analyse how ω_w varies with y (i.e., along the valley axis). The IGW frequency ω_w is determined by the dominant peak of the vertical velocity frequency spectrum computed at a location well above the katabatic flows. Note that, for an altitude lower than 500 m above the ground surface, because of the standing IGW system that develops and of possible wave-wave interactions, no clear frequency can be identified in the spectrum. Above 500 m by contrast, any altitude may be chosen to compute this spectrum; indeed, since the wave field propagates in a homogeneous medium, the wave frequency does not vary with z (as we checked it).

The frequency spectrum of the vertical velocity w was computed for $x = -0.6 \text{ km}$ at an altitude of $z = 2200 \text{ m}$ (i.e., above the height of the plateaus), for different values of y along the valley axis. We recall that the maximum slope angle of the topography α_{\max} in a given vertical (x, z) plane varies with y , and so does the frequency $\omega_k = N \sin \alpha_{\max}$ of the oscillations of the katabatic winds. The ratios ω_w/N and ω_k/N computed at this (x, z) location are plotted versus y in Fig 7a. The ratio ω_w/N appears to be independent of y and therefore of the slope angle. Hence, the IGW frequency does not seem to be imposed by the frequency ω_k of the oscillations of the katabatic winds. More precisely, ω_w varies as $C_w N$,

323 with C_w in the range 0.7 – 0.95. A representative value of C_w may be taken as its y -average
 324 value, equal to 0.82.

325 Note also that, for steeper slopes, the values of ω_w are closer to that of ω_k because ω_k
 326 increases with α_{\max} so that the absence of relation between ω_w and ω_k is more difficult to
 327 assess there.

328 4.3 Wavelength of the internal gravity wave field

329 One striking feature of the dispersion relation (1) is that only a time scale comes into play,
 330 namely $1/N$, but no length scale. Nevertheless, IGWs usually develop with a well-defined
 331 wavelength, which is imposed by an external length scale. This length scale may be fixed
 332 by the geometry of the forcing, as for lee waves, or from dimensions of the reservoir which
 333 contains the IGWs, as for seiches in lakes. In this section, we compute the wavelengths of
 334 the IGW field in all three directions and attempt to determine the external length scale which
 335 sets them.

336 The wavelength can be determined by plotting contours of the vertical velocity field
 337 (or potential temperature field) in a time-space diagram (also referred to as a Hovmöller
 338 diagram). For a monochromatic wave field, a (t, x_i) diagram provides the phase speed c_i
 339 in direction x_i , given by the slope of the contours, as well as the wave period T_w and the
 340 wavelength λ_i in that direction. We recall that $c_i = \lambda_i/T_w$, with $T_w = 2\pi/\omega_w$.

341 *Wavelength along the valley axis*

342

343 Contours of the vertical velocity field w in a (t, y) diagram are plotted in Fig. 8a for $x =$
 344 -0.6 km at a height of 800 m above the ground surface for the first 80 min of simulation.
 345 The IGWs reach this height at about 40 min into the simulation. The slope of the contours,
 346 equal to the phase speed c_y , is infinite for y larger than about 7 km (i.e., within the valley).
 347 Since the IGW period T_w is finite, this implies that the wavelength along the y direction
 348 is infinite for y larger than 7 km. Therefore, the IGW field may be assumed to be two-
 349 dimensional (i.e., the IGWs propagate in the (x, z) plane) beyond this distance.

350 Fig. 8a also clearly shows that IGWs are first emitted toward the valley end, along the
 351 longest and steepest slopes (for y larger than about 15 km), the wave emission progressively
 352 extending to the shallower slopes with time.

353 *Wavelength along the vertical*

354

355 Contours of the vertical velocity field w in a (t, z) diagram are plotted in Fig. 8b for
 356 $x = -0.6$ km and $y = 15$ km for the first 80 min of simulation. As above, phase lines
 357 are clearly visible, with a well defined slope c_z in the upper part of the phase lines. The
 358 value of this slope is equal to -2.5 m s^{-1} . The IGW period T_w is given by the distance
 359 along the horizontal axis between two maxima of vertical velocity; one finds $T_w \approx 10$ min,
 360 consistent with the value of $\omega_w \simeq 0.8N$ with $N = 0.0147$ rad s^{-1} . With $\lambda_z = c_z T_w$, one
 361 gets $\lambda_z \approx 1300$ m. Similar phase speed and period are found for any $y > 7$ km. Since the
 362 maximum height of the surrounding plateaus along the valley axis H is equal to 1700 m,
 363 one may conclude that λ_z is set by the valley depth.

364 *Wavelength along the cross-valley direction*

365 The wavelength along the cross-valley direction λ_x is simply inferred from $\tan \phi = \lambda_x / \lambda_z$.
 366 The relation $\omega_w \simeq 0.8N$ yields $\phi \simeq 37^\circ$ (since $\cos \phi = 0.8$) so that $\lambda_x \approx 0.75 \lambda_z \approx 1000$
 367 m. We note that λ_x is close to the width of the valley floor $2V_x = 1240$ m for the topography
 368 T1. As an alternate route to find the external length scale which controls the wavelength,
 369 one may conclude that λ_x is fixed by the width of the valley floor; λ_z would then result from
 370 the dispersion relation (1). However, Chemel et al. (2009) found that λ_x remains unchanged
 371 when the width of the valley floor is doubled, the height of the surrounding plateaus being
 372 unchanged. Hence, the results suggest that λ_z is set by the valley depth, λ_z imposing λ_x
 373 through the dispersion relation (1) with $\omega_w \simeq 0.8N$.
 374

375 **5 Sensitivity experiments**

376 In the previous section, we found that the IGW frequency ω_w is equal to $C_w N$, with C_w
 377 in the range $0.7 - 0.95$, and is independent of the frequency ω_k of the oscillations of the
 378 katabatic winds. The purpose of this section is to investigate the effects of the background
 379 stratification of the atmosphere and valley geometry on the characteristics of the IGW field.
 380 We first show that the simulations used for the sensitivity experiments can be run more
 381 efficiently by slightly changing the initial condition in the soil layers.

382 **5.1 Influence of the initial ground surface temperature**

383 In simulation S1, the initial temperature of the ground surface T_s was the same as those of
 384 the near-surface air T_a and deep soil T_2 (see Table 1). Our hypothesis is that, if T_2 is lower
 385 than T_a at the initial time, the katabatic flows should be established faster, and so the IGWs
 386 should develop more rapidly. In order to test this hypothesis, we initialized T_s and T_2 in
 387 simulation S1 such as $T_s - T_a = -3$ K and $T_2 - T_a = -5$ K (see simulation S5 in Table 1).
 388 By doing so, the ground surface temperature cools faster, and so does the near-surface air
 389 (not shown). As a result, the response of the atmosphere to the surface cooling is more rapid
 390 in simulation S5 than in simulation S1 but is qualitatively the same. More precisely, Fig. 9
 391 shows that results of simulation S1 corresponds to those of simulation S5 but delayed by 20
 392 min. To save computing time, all the simulations used for the sensitivity experiments were
 393 performed with the same initialization of the ground surface and deep soil temperatures as
 394 simulation S5 (see Table 1).

395 **5.2 Influence of the background stratification**

396 In order to investigate the effects of the background stratification of the atmosphere, we
 397 performed 8 simulations (simulations S2 to S9, see Table 1) with different values of the
 398 buoyancy frequency, ranging from 0.91×10^{-2} to 2.33×10^{-2} rad s⁻¹, corresponding to an
 399 initial stratification $d\theta/dz$ ranging from 2.3 to 15 K km⁻¹.

400 In agreement with the results of Sect. 4.2, the IGW frequency ω_w is found to be indepen-
 401 dent on y and therefore on the slope angle for every simulation. We shall therefore assume
 402 that the IGW frequency is nearly constant along y and compute its y -averaged value, $\langle \omega_w \rangle$,
 403 for each value of N , at a given (x, z) location (as before, for $x = -0.6$ km and $z = 2200$ m).

404 The ratio $\langle \omega_w \rangle / N$ is plotted in Fig. 10 for each value of N , in the simulations S2 to S9. Con-
 405 sistent with Sect. 4.2, $\langle \omega_w \rangle$ varies as $C_w N$, with C_w in the range 0.7 – 0.95. The figure also
 406 shows that C_w is approximately constant and about 0.8 for weak stratification ($d\theta/dz \leq 6$
 407 K km^{-1}) and slightly decreases with N when the stratification becomes stronger.

408 5.3 Influence of the topography

409 Coincidentally, for the topography T1 used in simulations S1 to S9, the maximum slope
 410 angle α_{\max} is about 45° at the valley end, and so $\sin \alpha_{\max} \approx 0.7$, which is close to the ratio
 411 $\omega_w / N \approx 0.8$ found hitherto for the IGW field. In order to remove any doubt concerning the
 412 non dependence of ω_w / N on α_{\max} (and thus on ω_k), a simulation was performed with the
 413 topography T2, for which the maximum slope angle is approximately constant, at a value of
 414 about 30° (see Sect. 2.2).

415 The characteristics of the IGW field are in line with the results of Sect. 4. The frequency
 416 spectrum associated with the time series of u_s at $x = -1.2$ km and $y = 15$ km at 12.5 m
 417 above the ground surface (see Fig. 11a) shows that the katabatic winds undergo temporal
 418 oscillations at a frequency $\omega_k = 6.4 \times 10^{-3} \text{ rad s}^{-1}$, giving $\omega_k / N \approx 0.5$ (namely $\sin 30^\circ$).
 419 The frequency spectrum associated with the time series of w at $x = -1.2$ km and $y =$
 420 15 km at 4000 m above the ground surface (see Fig. 11b) exhibits a peak at frequency
 421 $\omega_w = 11.8 \times 10^{-3} \text{ rad s}^{-1}$, giving $\omega_w / N \approx 0.8$. Thus, we can conclude with no doubt that
 422 the frequency ω_w of the IGWs is distinct from that of the oscillations of the katabatic winds
 423 ω_k and that the ratio ω_w / N is independent of the valley geometry.

424 6 Conclusions

425 The purpose of this work was to extend the study by Chemel et al. (2009) by investigating
 426 the effects of the background stratification and valley geometry on the characteristics of the
 427 IGW field generated by katabatic winds in a deep valley. For this purpose, a set of numerical
 428 simulations, for a wide range of stratifications and slope angles, were run for 3 hours during
 429 a winter night at mid-latitude and analysed.

430 The present study confirms that two oscillatory systems, spatially decoupled, coexist,
 431 consisting of (i) along-slope temporal oscillations of the katabatic winds and (ii) oscillations
 432 associated with the IGW field emitted by the katabatic winds, which propagates away from
 433 the slopes. The frequency of the oscillations of the katabatic winds ω_k is found to be equal
 434 to N times the sine of the maximum slope angle. The IGW frequency ω_w is found to be
 435 independent of α and about $0.8 N$.

436 The novelty here is to analyse the generation of the IGW field and the variations of the
 437 IGW frequency and wavelength in space, and as a function of the stratification and valley
 438 geometry.

439 The IGWs first form at the bottom of the slopes, as a result of a hydraulic jump in the
 440 katabatic flow, and are then emitted all along the slope. The IGWs are generated along the
 441 longest and steepest slopes (i.e, as one moves toward the valley end) and, later in time, along
 442 shorter and shallower slopes as well. The IGW field propagates in a plane perpendicular
 443 to the valley axis and is therefore two-dimensional. Whatever the location in the valley
 444 atmosphere (away from the katabatic flow), its frequency ω_w varies as $C_w N$, with C_w in the
 445 range 0.7 – 0.95. The simulations used for the sensitivity experiments with different values
 446 of background stratification indicated that the ratio ω_w / N is constant and about 0.8 for weak

447 stratification ($d\theta/dz \leq 6 \text{ K km}^{-1}$) and slightly decreases with N when the stratification
 448 becomes stronger. The present analysis also showed that the ratio ω_w/N is independent of
 449 the valley geometry and that the IGW wavelength is controlled by the valley depth.

450 The fact that the ratio ω_w/N varies in a narrow range of values around 0.8 may be
 451 explained by the theoretical work of Voisin et al. (2011) (see also Voisin (2007)). This work
 452 shows that the power of the IGWs radiated by an oscillating sphere or cylinder displays a
 453 maximum value for an oscillating frequency close to $0.8 N$ (the radiated power is the wave
 454 energy averaged over the period of the oscillating source, divided by the period). When an
 455 unsteady flow, such as a katabatic flow, is considered instead of an oscillating sphere, one
 456 may argue that this flow possesses a large range of frequencies, among them those with a
 457 frequency close to $0.8 N$ are the most powerful at emitting IGWs, and therefore dominate the
 458 IGW signal. Laboratory experiments of localized turbulence in a stably stratified fluid are
 459 consistent with this result, reporting that the IGWs propagate at a fixed angle with respect
 460 to the horizontal, of about 45° (e.g. Wu 1969; Cerasoli 1978; Dohan and Sutherland 2003;
 461 Taylor and Sarkar 2007).

462 **Acknowledgements** The PhD grant of YL was funded by the French Région Rhône-Alpes as part of the
 463 ‘Cluster Environnement’. YL and CS also acknowledge financial support from the French ‘Institut national
 464 des sciences de l’Univers’ (INSU) as part of the LEFE/IDAO program. Time-consuming computations were
 465 performed thanks to the French IDRIS national supercomputing facilities.

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Tables

Table 1 Description of the simulations. T_s and T_a are the temperature of the ground surface and the near-surface air, respectively, and T_2 is the deep soil temperature; N is the buoyancy frequency. All values are initial values.

Simulation	$T_s - T_a$ (K)	$T_2 - T_a$ (K)	$d\theta/dz$ (K km ⁻¹)	N (rad s ⁻¹)	Topography
S1	0	0	6.0	1.47×10^{-2}	T1
S2	-3	-5	2.3	0.91×10^{-2}	T1
S3	-3	-5	3.4	1.11×10^{-2}	T1
S4	-3	-5	4.7	1.30×10^{-2}	T1
S5	-3	-5	6.0	1.47×10^{-2}	T1
S6	-3	-5	8.0	1.70×10^{-2}	T1
S7	-3	-5	10.0	1.90×10^{-2}	T1
S8	-3	-5	12.3	2.11×10^{-2}	T1
S9	-3	-5	15.0	2.33×10^{-2}	T1
S10	-3	-5	6.0	1.47×10^{-2}	T2

Figures

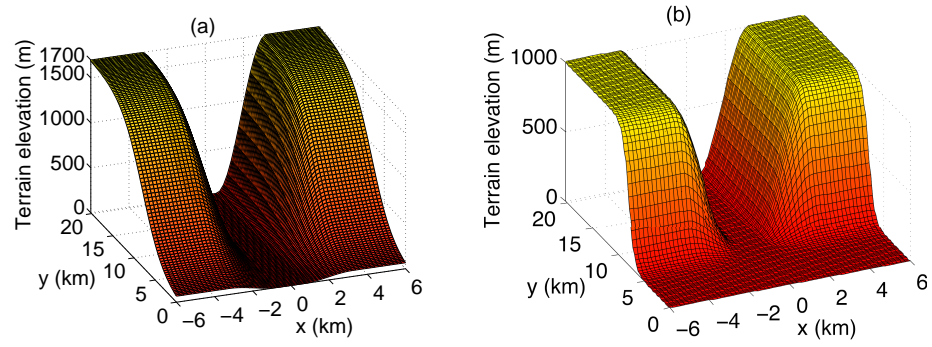


Fig. 1 Topography of the valleys (a) T1 and (b) T2 (see Section 2.2 for details). Note that the valley floor is flat

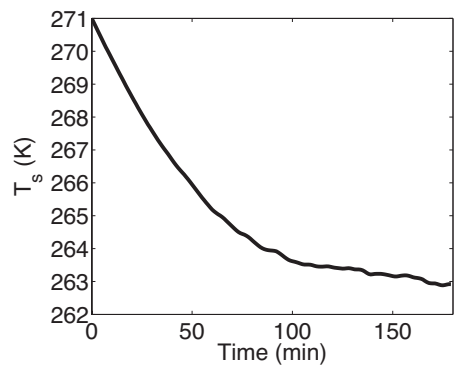


Fig. 2 Time series of the ground surface temperature T_s half way down the slope at $y = 15$ km for simulation S1

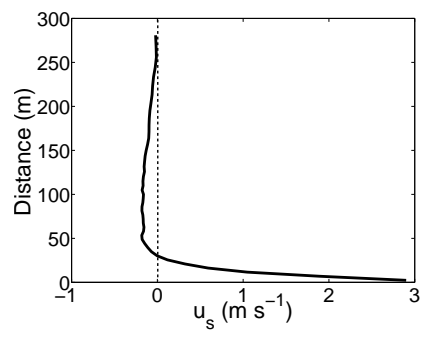


Fig. 3 Along-slope component u_s of the katabatic wind versus distance along the vector normal to the sloping surface, for $y = 7$ km and $t = 74$ min, at the x location of maximum slope angle (i.e., half way down the slope), for simulation S1

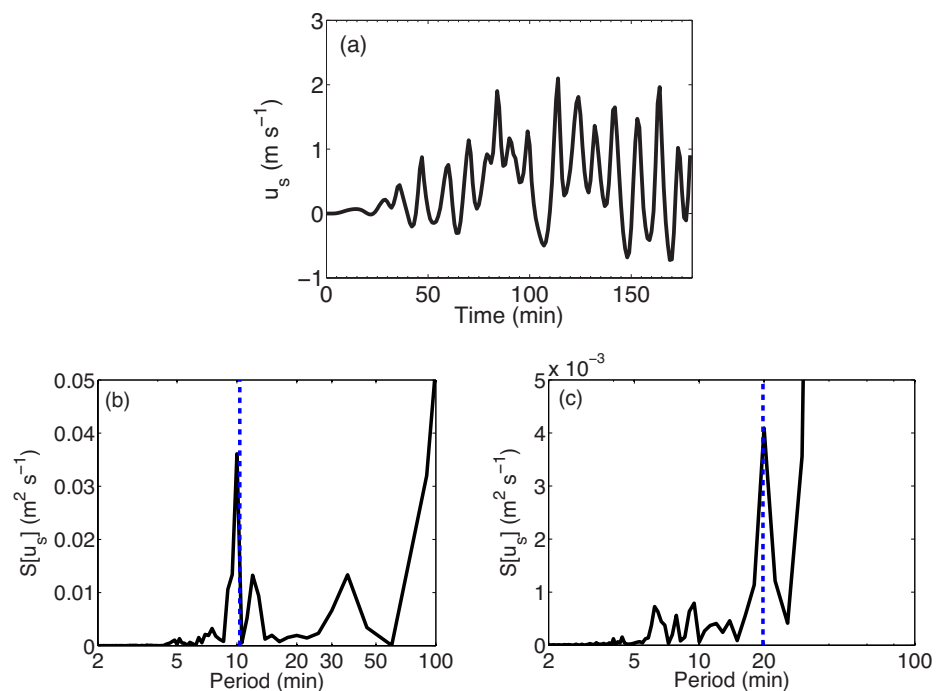


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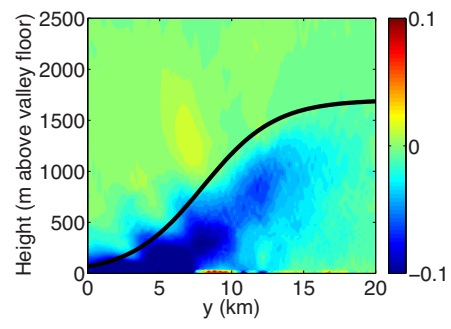


Fig. 5 For simulation S1: colour-filled contours of the along-valley component v of the wind, in m s^{-1} , in a (y, z) vertical cross section for $x = 0$ km and $t = 45$ min. The black solid line indicates the height of the plateaus along the valley axis

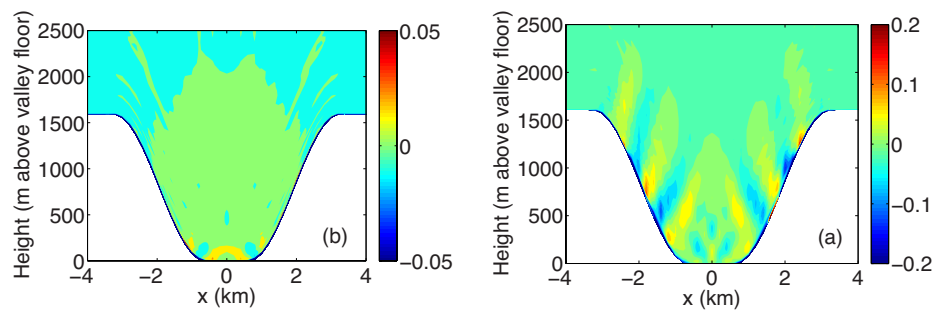


Fig. 6 For simulation S1: colour-filled contours of the vertical velocity component w , in m s^{-1} , in a vertical cross section for $y = 15$ km at **(a)** $t = 25$ min and **(b)** $t = 45$ min

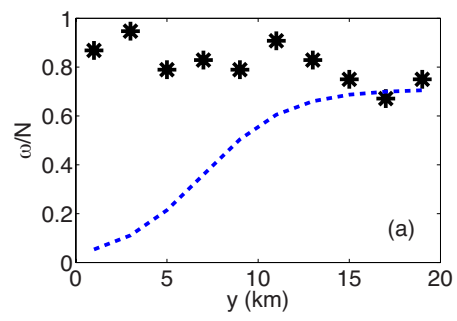


Fig. 7 For simulation S1: ratio of the IGW frequency ω_w to the buoyancy frequency N (*), and of the frequency of the oscillations of the katabatic wind ω_k to N (blue dashed line) versus y (i.e., along the valley axis) for $x = -0.6$ km at an altitude of 2200 m

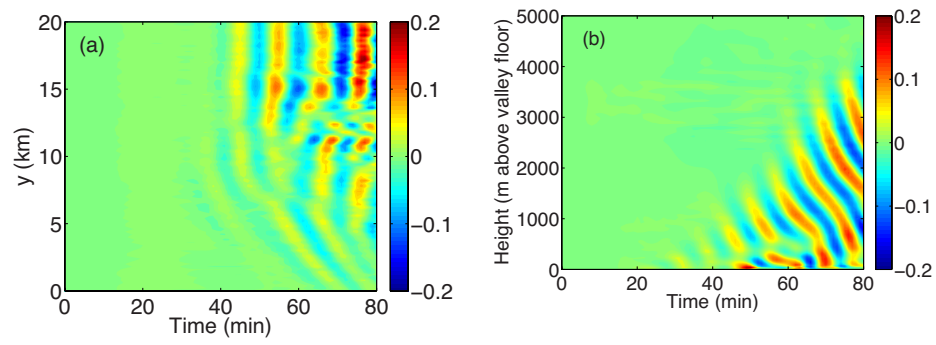


Fig. 8 For simulation S1: colour-filled contours of the vertical velocity component w , in m s^{-1} , (a) in a (t, y) diagram for $x = -0.6$ km at a height of 800 m above the valley floor and (b) in a (t, z) diagram for $x = -0.6$ km and $y = 15$ km

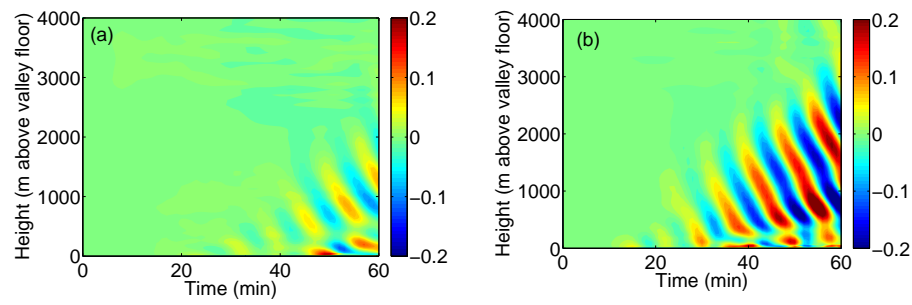


Fig. 9 Colour-filled contours of the vertical velocity component w , in m s^{-1} , in a (t, z) diagram for $x = -0.6$ km and $y = 15$ km **(a)** for simulation S1 and **(b)** for simulation S5

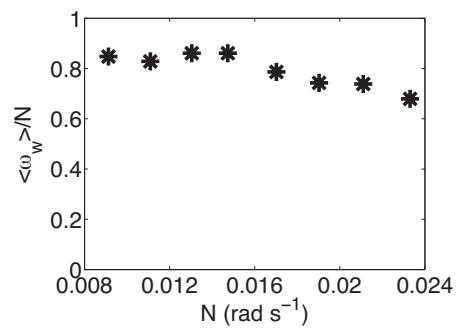


Fig. 10 Ratio of the mean value of the IGW frequency ω_w over y (i.e., along the valley axis), denoted by $\langle \omega_w \rangle$, and the buoyancy frequency N (*), for $x = -0.6$ km at an altitude of 2200 m, for each value of N in the simulations S2 to S9

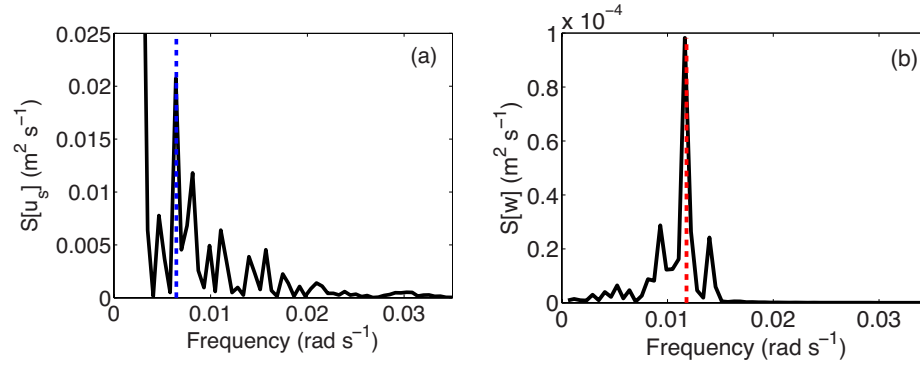


Fig. 11 For simulation S10: **(a)** Frequency velocity spectrum $S[u_s]$ of the along-slope component of the wind u_s at $x = -1.2$ km and $y = 15$ km at 12.5 m above the ground surface; **(b)** Frequency velocity spectrum $S[w]$ of the vertical component of the wind w at $x = -1.2$ km and $y = 15$ km at 4000 m above the ground surface. The blue dashed line in plot **(a)** indicates the frequency of the oscillations of the katabatic winds ω_k calculated using Equation 5. The red dashed line in plot **(b)** indicates the IGW frequency $\omega_w \approx 0.8N$