Circular Polarization in Scattered Light as a Possible Biomarker

W. B. Sparks^{a*} J. H. Hough^b, L. Kolokolova^c, T. Germer^d, F. Chen^e, S. DasSarma^e, P. DasSarma^e, F. Robb^e, N. Manset^f, I.N. Reid^a F. Duccio Macchetto^a William Martin^b

a- Space Telescope Science Institute, Baltimore

b- Center for Astrophysics Research, Univ. of Hertfordshire

c- Dept. of Astronomy, Univ. of Maryland

d- National Institute of Standards and Technology

e- Center for Marine Biotechnology, Univ. of Maryland Biotechnology Institute f- Canada-France-Hawaii Telescope

* Corresponding author. Tel. (1)410-338-4843 *E-mail address*: sparks@stsci.edu (William B. Sparks)

Pages: 17

Tables: 0

Figures: 4

Abstract

Biological molecules are optically active and exhibit homochirality. Therefore, it is possible that the scattering of light by biological molecules might result in a macroscopic signature in the form of circular polarization. If this is the case, then circular polarization spectroscopy, which may be utilized in remote sensing, can offer a powerful indicator of the presence of a universal biosignature, namely homochirality. Here, we describe laboratory experiments designed to investigate this idea. We focus on photosynthetic microorganisms, and also show results from macroscopic vegetation and control minerals. In the microorganisms, we find unambiguous circular polarization associated with electronic absorption bands of the photosynthetic apparatus. Macroscopic vegetation yields a stronger and more complex signature while the control minerals produce low-levels of circular polarization unrelated to their spectra. We propose a heuristic explanation of our results, which is that the polarization is produced by absorption-band circular dichroism in the layer between optical depth zero and one. The results are encouraging for the use of circular polarization spectroscopy as a remote sensing biomarker.

Key words:

. Introduction

All known living material exhibit a remarkable property which is homochirality – the complex chiral molecules of biology occur almost exclusively with only a single handedness (chiral: from Greek, $\chi\epsilon(\rho, cheir, hand)$). For example, the vast majority of organisms use only left-handed L-amino acids in proteins and right-handed D-sugars in nucleic acids. More generally, a chiral compound contains one or more asymmetric centers and thus can occur in non-superimposable mirror-image forms. This chiral property is believed to be a necessity for self-replication, and as a consequence, it is likely to be a generic property of all biochemical life whether similar to terrestrial or not.

Biological material displays optical activity, arising from either circular dichroism (different absorption coefficients for left and right circularly polarized light) or differential scattering of left and right circularly polarized light by its component chiral molecules [1-3]. Circular dichroism spectroscopy is a standard analysis technique for studying protein structure [4,5]. For example, in this regard, chlorophyll has significantly circularly dichroic absorbance bands which leads to a distinctive dimer circular polarization signature [6]. Circular Intensity Differential Scattering (CIDS) considers the circular polarization induced by scattering from chiral samples and has been studied by [7]. Circular polarization can also be caused by optical interaction associated with the birefringence of macroscopic biopolymer membranes of a biological structure; while light emitted through fluorescent processes, commonly associated with molecules of biological significance, may also exhibit circular polarization. In photosynthesis (discussed below) circular dichroism and related phenomena are induced by the intrinsic chirality of the molecules, from excitonic coupling between chromophores in molecular complexes and from macroscopic organization of the system [8]. A variety of microbes were found to produce fractional polarizations $\sim 10^{-2}$ to 10^{-3} [9]. Here, we are interested in the fact that interaction between light and living organisms can produce a detectable circular polarization signature by means of several mechanisms, and hence that circular polarization spectroscopy may be useful as a remote sensing indicator of the presence of biological material.

Abiotic contributions to circular polarization arise from atmospheric and aerosol scattering and from mineralogical scattering processes. Some minerals and crystals are also optically active, however integrating over a naturally produced sample is expected to give equal fractions of enantiomorphs so that the optical activity of the macroscopic sample averages to zero. Multiple reflection and phase effects can also introduce circular polarization; however, these typically produce a smooth distribution recognizably related to the geometry of the scattering [10, 11], and scattering processes and mineralogical optical activity also typically have a very broad, smooth spectral dependence. This is in contrast to electronic transitions associated with biological material which present relatively narrow absorption bands with which the strongest circular polarization signal is associated (the Cotton effect). Even if local enantiomeric excesses of optically active minerals were to exist, it is expected that they would be distinguishable from biological matter by their spectral properties [1, 5]. Empirically, Solar System measurements of circular polarization consistently find low integrated fractional polarization levels, ~10⁻⁴ to 10⁻⁵ [10, 12, 13], some two or three orders of magnitude smaller than possible biological effects.

An especially promising type of interaction between light and living organisms is photosynthesis and phototrophy ("light harvesting"; hereafter we loosly use photsynthesis to include phototrophy). In photosynthesis (and phototrophy), the organism extracts energy by triggering an electronic excitation of the active pigment in diffuse absorbance bands.

From an astronomical perspective, photosynthetic activity is extremely attractive, since it is largely a *surface phenomenon*, which is required for external optical observation. Photosynthesis generally requires interaction with light of the host star, hence is tuned to the *wavelength of*

maximum flux and is further tuned to the wavelength of maximum transmission of the planetary atmosphere [14]. These conditions optimize the possibility of observing photosynthesis.

Biophysically, the strength of the interaction between organisms and light is particularly strong in photosynthesis, and the strongest signatures of chirality are manifested in the circular dichroism of the electronic absorption bands the characterize photosynthetic activity [8]. Hence, we by focusing on photosynthetic processes we optimize our chances of finding a circular polarization signature.

Photosynthesis offers enormous evolutionary advantages and arose early in the history of the Earth. Once present, over 3 Gyr in the past, photosynthesis by oceanic microbial life dramatically changed the atmosphere of the Earth, giving rise to oxygen in large quantity, and has played a fundamental role in life's progression ever since. Hence photosynthesis, oxygenic or otherwise, is likely to be a common and successful life strategy that extraterrestrial organisms could certainly utilize. Therefore, a microbial dominated photosynthetic biosphere is one of the most likely possibilities for a randomly chosen extrasolar planet.

For these reasons, Sparks et al. (2009) [15] studied the circular polarization spectra of scattered and transmitted light from photosynthetic cyanobacteria, which were found to produce significant circular polarization spectral signature associated with the photosynthetic pigments. They also presented results from the bacteriochlorophyll-based purple non-sulphur bacteria *R. rubrum*, and an example of a leaf and mineral for control. Here, we present new experimental results for the cyanobacteria, specifically comparing low and high optical depths, we describe initial attempts to measure halobacteria, more macroscopic vegetation results and present additional mineral control samples.

2. Experimental Concept and Choice of Samples

Our aim is to understand whether astrobiologically important microbes produce a macroscopic circular polarization signature in scattered light. To address this we established a

dedicated laboratory configuration, described in detail below, at the National Institute of Standards and Technology (NIST) and obtained samples of microbes in liquid suspension from the Center of Marine Biotechnology, part of the University of Maryland Biotechnology Institute (COMB).

Our primary microbial targets are marine cyanobacteria, which are photosynthetic prokaryotes possessing chlorophyll a and their antenna pigments phycocyanin and phycoerythrin. They have an extensive fossil record that may date back 3.5 billion years [16]. With the early development of a photosynthetic capability, they are widely thought to have been the primary agent responsible for the rise of oxygen in the primitive atmosphere on Earth. Cyanobacteria are abundant and ubiquitous in marine and freshwater systems; are found in cold and hot environments; tolerate conditions of high desiccation; and can produce biochemical products that provide protection from UV damage. High diversity and great adaptability have made cyanobacteria extremely competitive on the planet Earth. These species have been studied extensively as part of exobiological research on the limits of life in the Solar system [17] and are now being studied as an analogue for past (or perhaps even present) life on Mars. Endolithic communities (living embedded in the surface of rocks) of photosynthetic cyanobacteria have been found in subsurface layers in the Mars analogue arctic Svalbard rocks and elsewhere, suggesting additional strategies for life that may find extraterrestrial application. Endolithic colonies of hardy, photosynthetic cyanobacteria have also been found just mm below the surface in halite evaporite rocks from the most hostile, hyperarid parts of the Atacama desert [18]. This extremely dry, saline, evaporite environment is also a plausible terrestrial Mars analogue site [19]. By residing just below the surface of the rock, light useful for photosynthesis reaches the organisms, while potentially harmful UV radiation does not.

We also initiated a study of phototrophic haloarchaea, which are salt-loving extremophiles, robust against UV radiation and tolerant of extreme desiccation. Haloarchaea are commonly

found in salt-rich lake environments, and imbedded in evaporative products such as halite associated with those environments. In recent experiments they have been shown to possess exceptional hardiness to Solar UV and electron beam radiation. This hardiness and the tolerance of haloarchaea to harsh cold, dry, salty conditions are qualities that a microbial species alive on Mars either now or during a moist past would need to share [20, 21]. Further, they use a light harvesting system that contains the retinal-based pigment bacteriorhodopsin in their metabolism [22].

For context, we present measurements of examples of macroscopic vegetation chosen largely by availability near the laboratory, and control minerals sulphur (which has a strong spectral "edge") and a Mars regolith analogue, JSC-1.

3. Experimental Methods

Fig. 1

We purchased a sophisticated dedicated polarimeter optimized for measurement of circular polarization in the presence of significant linear polarization. Our requirement was that a circular polarization degree of 10⁻⁴ (i.e. polarization percent P=0.01%) be measurable in the presence of linear polarization of a few percent and that the polarimeter be tunable from 400 nm to 800 nm, essentially the entire visual spectrum with a modest extension into the near infrared. The long wavelength cutoff was designed to extend beyond chlorophyll's red edge at 700 nm while covering the important absorption bands in the optical region. The equipment is hosted at NIST in the optical metrology department. It has been well-characterized and found to meet its design requirements.

Details of the experiment may be found in [15]. The polarimeter is a Hinds Instruments, Series II/FS42-47 dual photoelastic modulator (PEM) precision optical polarimeter. The target performance figure of achieving a polarization measurement of 1:10⁴ for circular polarization level (i.e. 0.01%) in the presence of linear polarization degree of 0.03 (i.e. 3% linear

polarization) is achieved. The polarimeter is oriented vertically with a direct unobscured view of the sample. Our reflection experiments utilize an integrating sphere which depolarizes and diffuses the illuminating light and provides a distribution of incidence angles onto the sample. The source of illumination can also be targeted onto a white plaque below the sample, and light passes through the sample into the polarimeter, yielding a transmission spectropolarization measurement to complement the reflectance experiments. The transmission configuration measurement is analogous to a classical circular dichroism measurement. The polarimeter and experimental configuration is shown in Fig. 1.

4. Experimental Results

4.1 Cyanobacteria

Detail descriptions of the sample preparation techniques may be found in [15]. One of the specimen results presented therein was to compare the transmission and absorption polarization spectra for the organism *Synechococcus* WH8101 which appears green due to the presence of the light harvesting pigment phycocyanin. There, it was shown that the strongest circular polarization features in both transmission and absorption are related to the electronic absorption bands of the photosynthetic pigments. In order to obtain an initial insight into the influence of microbe concentration on polarization levels, we carried out an additional experiment in which a very much more concentrated sample of the same organisms was measured in the reflection configuration. To the eye, the appearance was of a very dark green, almost black. Numerically, the reflected light intensity of the concentrated sample is 10 to 30× fainter than the "shallow" sample presented in [15]. Fig. 2 A summarizes the polarization results for the shallow reflection and transmission experiments for WH8101, as described in the caption. Fig. 2 B shows the data for the "deep" or concentrated dark sample. Overlaid for comparison are the polarization curves for both the shallow reflected experiment and the transmission experiment. It can be seen firstly

Fig. 2

that the shallow reflection and transmission measurements result in essentially the same degree of polarization. There are prominent features at the location of chlorophyll $a \approx 680$ nm), phycocyanin (≈ 620 nm) and a mix of carotenoids in the blue. Each feature is characterized by both a relatively well-defined absorption band, and a significant polarization within the band. We interpret this as the Cotton Effect, where the relatively strong circular dichroism of the electronic transitions is manifested in the transmission and reflection spectra as circular polarization. For the chlorophyll-a band, the polarization switches sign at the location of the absorption maximum, evidence of the excitonic dimer transition.

The concentrated sample has a similar but not identical spectrum. The phycocyanin absorption band is no longer evident, and the blue wing of the chlorophyll absorption is reduced, though the red carotenoid absorption is slightly higher. This can be understood as adaptation to the strong illuminating light needed to acquire a measurable polarization signal. (These data were smoothed more heavily in the spectral dimension to increase signal to noise per point). The phycocyanin pigment either saturated or the organism suppressed its production. Apart from these changes, the character of the polarization spectrum is similar, with features at the absorption features and an overall quantitative polarization level that is very close to the polarization degree returned by the shallow sample. We propose, below, that this is due to an optical depth effect, whereby circular polarization is produced in the circularly dichroic layer between optical depths zero and one.

4.2 Halobacteria

In similar fashion, we obtained spectropolarization measurements of halobacteria in suspension, strains S9 and NRC-1. These experiments revealed absorption spectra that did not contain prominent relatively narrow bands characteristic of the cyanobacteria photosynthesis. Instead, the spectra were smooth with only a small amplitude, very broad feature at the location of the anticipated light harvesting pigment bacteriorhodopsin 550 nm absorption maximum. The

presence of gas vesicles within the microorganisms and possible additional scattering substances in the suspension resulted in a smooth spectrum characteristic of a broad band scattering phenomenon, rather than the light harvesting absorption spectra. Overall circular polarization levels either were low (less than 10⁻⁴) or were offset globally from zero as might be expected from a scattering scenario. In some cases, there were hints of structure in the polarization spectrum, however these were at a very low level. More work is needed to understand the situations in which the organism produces a light-harvesting dominated spectrum, and hence we may at that stage predict the expected polarization levels from the circular dichroism properties of bacteriorhodopsin.

4.3 Vegetation

Fig. 3

For context, we also acquired polarization spectra of a variety of easily obtained macroscopic vegetation from the NIST grounds. An example of a maple leaf spectrum was presented in [15]. The average circular polarization reflection spectrum shown in Fig. 3 is an average of scans of three maple leaves, three oak leaves, two clover leaves and one leaf from a cherry tree. The reflectance spectra are all very similar, the cherry tree leaf shown as an example. The circular polarization reflection spectra however do show substantial diversity. The average spectrum brings out the common features, which are a strong polarization signal associated with chlorophyll absorption between 650 and 700 nm, and non-zero polarization with a broader spectral dependence to the blue. Polarization is low at wavelengths longer than the cholorphyll red edge at 700nm. In a planetary context, it will be of interest to establish the polarization characteristics of large scale, sometimes heterogeneous, vegetation dominated scenery such as forests or grasslands. In addition to the mix of scattering vegetation, there is a mix of scattering angles and the potential for multiple scattering within the forest canopy, or between closely packed grasses. Empirical measurements are needed to determine whether a circular polarization

signal is present, and the physical processes by which it is produced, and whether in turn it may provide a useful diagnostic observable.

To guard against possible instrumental polarization artefacts arising from strong spectral

4.4 Mineral Controls

features, we made measurements of two minerals exhibiting significant strength "edges" in their spectra. The results from the analysis of one of these, iron oxide, were presented in [15]. The other was from a naturally occurring sulphur rock from Hawaii, dimensions two to three inches. The rock appears to be essentially pure sulphur, with some coarse sulphur crystalline structures imbedded in the matrix. Fig. 4 presents the polarization spectrum derived from a series of measurements of this rock. While there are small non-zero polarization levels evident, they do not share the distinctive character of relationship to the intensity spectra that the photosynthetic cyanobacteria or leaves do. The character of this mineral is quite different to that of the biological specimens. Additional polarization spectra of a variety of other minerals were presented by [1] and this was true in all cases.

Finally, we obtained spectropolarimetry of a popular Mars regolith analogue. Martian regolith simulant JSC Mars-1 is made of weathered volcanis ash collected from a cinder cone of the Island of Hawaii [23, 24]. The collected tephra was dried and passed through 1 mm sieves to keep only the finer material. This simulant has spectral similarities to the Martian bright regions – the VIS spectrum contains a relatively featureless ferric absorption edge and an indication of a ferric absorption band in the 800-1000 nm region. To the eye, the material is a powder which appears fine grained, smooth and grey. The spectrum obtained with the Hinds polarimeter is also smooth with essentially no structure across the visible wavelength region to which we have access. The circular polarization level is very close to zero throughout. There are small non-zero ripples in the polarization spectrum which may be due to multiple reflections and interference within the powder, however, as for the sulphur, there are no correlations with spectral features.

Fig. 4

5. Discussion

Our series of experiments was initiated to test the notion that circular polarization spectroscopy could provide a powerful biomarker, being directly sensitive to the chirality of the scattering medium. Homochirality presents an essentially unique biosignature, while circular polarization spectroscopy is amenable to standard astronomical or remote sensing applications. Hence it may be envisaged that major surveys of the surfaces of astrobiological targets including Mars, Europa and Titan can be undertaken.

We found that the polarization signal in reflection was almost the same for a very dark suspension as it was for a light, shallow one. Our conjecture is that this approximate constancy of polarization signal in reflected light is due to an optical depth effect. The optical depth $\tau = (\sigma_k + \sigma_s)n.d$, where d is depth, n the number density, and σ_k , σ_s are scattering and absorption cross-sections. If τ is dominated by bulk scattering (or absorption) which is relatively insensitive to circular polarization, then light undergoes multiple scattering beyond an optical depth of one, and such interior multiple scattering destroys any polarization information. Viewed from the outside, the "optical depth equals one" surface may therefore be considered as an unpolarized surface from which light emerges and then passes through the layer which is circularly dichroic within the absorption bands between optical depths one and zero (the surface of the medium). In the case of shallow reflection, for example, this unpolarized surface is provided by the reflection from the white spectralon plaque beneath the sample; in the concentrated case, by the internal scattering and absorption characteristics of the medium. This heuristic scenario requires detailed modeling to properly understand the interplay between scattering, absorption and resonant electronic absorption spectral features. Concerning application as a biomarker, if valid, the relevant parameter is optical depth rather than physical depth. So for example in deep clear water, a low concentration of microbes could yield approximately the same polarization signal as

a more concentrated suspension provided that the optical depth remains dominated by the microbes.

Of course in the Solar System, there are no other terrestrial-like oceans visible on the surface. Europa and other moons of Jupiter may have subsurface oceans, and Titan may have surface lakes of hydrocarbons that could potentially be accessible. Europa is one of the most important astrobiological targets in the Solar System. While Mars may have hosted liquid water oceans in the geological past, it is likely that Europa has a saline liquid water ocean at the present time [25]. Europa may therefore be able to sustain life. It is also thought that transport occurs between the interior ocean and the surface, from morphological studies of terrain features most reasonably explained by sub-glacial activity and geological youth of the surface [26, 27]. There are complex areas of brownish-red material that have apparently seeped onto the surface from the ocean below whenever the ice shell is fractured, punctured by impact or melted by internal heating, Fig. 3.2. This reddish material appears to be a complex mixture of sulfate hydrates and other materials [28]. [29] even considered the possibility of a microbial origin for features seen in the Galileo NIMS spectra. A detailed surface spectroscopic and polarimetric analysis of this previously interior material could reveal important information on its origin and constituents and represents an approach that should precede potential landers and ice penetrators.

Cyanobacteria and other photosynthetic microbes also exist in solid surface communities, and even in endolithic communities within rocks. Endolithic communities (living embedded in the surface of rocks) of photosynthetic cyanobacteria have been found in subsurface layers in the Mars analogue arctic Svalbard rocks and elsewhere, suggesting additional strategies for life that may find application on Mars. Endolithic colonies of hardy, photosynthetic cyanobacteria have also been found just mm below the surface in halite evaporite rocks from the most hostile, hyperarid parts of the Atacama desert [18]. This extremely dry, saline, evaporite environment is also a plausible terrestrial Mars analogue site[19]. By residing just below the surface of the rock,

light useful for photosynthesis reaches the organisms, while potentially harmful UV radiation does not. It would be extremely interesting to know whether a distinctive polarization signature survives. There is likely to be additional scattering from the surface roughness, and considerations of optical depth are replaced by the surface coverage of microbes, or equivalently their dilution by unobscured rock. This is work for the future.

Life at the surface of Mars has long been considered unlikely because of the lack of water, the high levels of ultraviolet radiation, and the lack of organics as determined from Viking soil analyses [30, 31]. On Mars, however, there are surface features that indicate the presence of considerable flows of water in the past, and recent results from the Mars Express and other missions show that substantial quantities of water and perhaps even methane are present [32, 33, 34]. These considerations, when taken together with the tolerance of halophilic bacteria (salt-loving and often found in evaporative conditions) to extreme environments and the possibility of organisms being preserved in the martian permafrosts [20,35], enhance the probability that life may be present at some surface locations, or may have been present in the past. If so, the timescale for racemization has been estimated to exceed the lifetime of the planet and a fossil record of chirality may exist at the present time [36] Recently, [37] revived the notion that the chemical changes seen by the Viking lander experiments may have a biological origin. Therefore, it would seem prudent to carry out a survey of the Martian surface to determine whether or not there are any locations that display significant circular polarization, which in turn could be an indication of past or present chirality in the surface material.

6. Conclusions

We have presented new evidence that microbial photosynthetic organisms can produce a measurable circular polarization signature in their spectra. The primary polarization signatures arise in concert with electronic spectral absorption bands, as in the Cotton Effect. Hence, for utility as a biomarker, it is essential that the polarization spectrum be viewed together with a

critical examination of the intensity spectrum. We showed that the strength of the polarization signal did not change substantially even with an increase of over a factor ten in the microbial concentration. We interpreted this as an optical depth effect and predict that in liquid suspension, the polarization signal in reflected light will be almost constant and equivalent to that for optical depth unity. Control minerals exhibiting spectral edges in their intensity distribution did not yield a spectropolarization signature, and the Mars regolith stimulant studied did not yield a significant spectropolarization signature. Macroscopic vegetation did provide strong polarizations. Not only may circular polarization spectroscopy measurements be made remotely, but also they depend for the presence of a significant signature upon a characteristic of life which is thought to be exceedingly general and a likely property of all biochemical life, namely its homochirality. In a variety of situations therefore, it appears the circular polarization spectroscopy could provide a powerful biomarker.

Acknowledgement

References

- [1] Pospergelis M.M., Spectroscopic measurements of the four Stokes parameters for light scattered by natural objects, Soviet Physics Astronomy, 1969: 12, 973-977
- [2] Wolstencroft R.D., The circular polarization of light reflected from certain optically active surfaces, 1974, Planets, Stars, and Nebulae: Studies with photopolarimetry, ed. Gehrels, University of Arizona Press, 495-499.
- [3] Wolstencroft R.D., Tranter G., Le Pevelen D.D., Diffuse reflectance circular dichroism for the detection of molecular chirality: an application in remote sensing of flora, 2002b, Bioastronomy 2002: Life among the stars, IAU Symp. 213, eds. R. Norris & F. Stootman, ASP, 149-153.
- [4] Purdie N., Brittain H.G., Analytical applications of circular dichroism, 1993, Techniques & instrumentation in analytical chemistry, 14, Elsevier.
- [5] Kelly S.M., Price N.C., The use of circular dichroism in the investigation of protein structure and function, Current Protein and Peptide Science, 2000:Vol. 1, 349-384.
- [6] Houssier, C., Sauer, K., Circular dichroism and magnetic circular dichroism of the chlorophyll and protochlorophyll pigments. *J Amer Chem Soc*, 1970: **92:4**, 779-791.
- [7] Bustamante, C., Maestre, M.F & Kelle, D., Expressions for the interpretation of circular intensity differential scattering of chiral aggregates. *Biopolymers*, 1985: **24** (8): 1595-1612.

- [8] Garab, G, Linear and circular dichroism. *Biophysical Techniques in Photosynthesis*, *Advances in Photosynthesis*, 1996: **3**, eds. J. Amesz & A.J. Hoff, p.11-40. Kluwer.
- [9] Salzman, G.C. & Gregg, C.T., Current and experimental methods of rapid microbial identification . *Biotechnology*, 1984: **2**, 243-248.
- [10] Kemp, J.C., Wolstencroft, R.D., Swedlund, J.B. Circular polarization: Jupiter and other planets. *Nature*, 1971: **232**, 165-168.
- [11] Bandermann, L.W., Kemp, J.C., Wolstencroft, R.D., Circular polarization of light scattered from rough surfaces . *MNRAS*, 1972: **158**, 291.
- [12] Swedlund, J.B., Kemp, J.C., Wolstencroft, R.D. Circular polarization of Saturn. ApJ, 1973: 178, 257-266.
- [13] Meierhenrich, U.J., Thiemann, W.H.-P., Barbier, B., Brack, A., Alcaraz, C., Nahon, L., Wolstencroft, R., Circular polarization of light by planet Mercury and enantiomorphism of its surface minerals. *Origins of Life and Evolution of the Biosphere*, 2002: **32**, 181-190.
- [14] Wolstencroft, R.D. & Raven, J.A., Photosynthesis: likelihood of occurrence and possibility of detection on Earth-like planets. *Icarus*, 2002: **157**, 535-548.
- [15] Sparks, W.B., Hough, J.H., Germer, T.A., Chen, F., DasSarma, S., DasSarma, P., Robb, F.T., Manset, N., Kolokolova, L., Reid, I.N., Macchetto, D., Martin, W. Detection of Circular Polarization in Light Scattered from Photosynthetic Microbes, PNAS 2009: submitted..
- [16] Whitton, B. A. and M. Potts. 2000. Introduction to the Cyanobacteria. pp. 1-11. In: The Ecology of Cyanobacteria: Their Diversity in Time and Space. Ed. B. A. Whitton and M. Potts. Kluwer Academic Publishers, Dordrecht, The Netherlands.
- [17] Wynn-Williams, D. D.. 2000. Cyanobacteria in Dessert Life at the Limit? pp. 341-366. In: The Ecology of Cyanobacteria: Their Diversity in Time and Space. Ed. B. A. Whitton and M. Potts. Kluwer Academic Publishers, Dordrecht, The Netherlands.
- [18] Wierzchos, J., Ascaso, C., McKay, C.P., Endolithic cyanobacteria in halite rocks from the hyperarid core of the Atacama desert, Astrobiology, 6, 415-422 (2006).
- [19] Navarro-González, R., Rainey, F.A., Molina, P., Bagaley, D.R., Hollen, B.J., de la Rosa, J., Small, A.M., Quinn, R.C., Grunthaner, F.J., Cáceres, L.,Gomez-Silva, B., McKay, C.P., Mars-Like Soils in the Atacama Desert, Chile, and the Dry Limit of Microbial Life, *Science*, 302, 1018-1021 (2003).
- [20]Landis G.A., Martian Water: Are there extant halobacteria on Mars?, 2001, Astrobiology, Vol. 1, No. 2, 161-164
- [21] DasSarma, S. Extreme Halophiles Are Models for Astrobiology Microbe, Volume 1, Number 3, p.120-127 (2006).
- [22] DasSarma, S., Extreme microbes, American Scientist, 95, #3, May-June 2007 (cover feature) (2007).
- [23] Carlton C. Allen, Richard V. Morris, et al. 1998, Lunar and Planetary Science XXIX
- [24] Carlton C. Allen, Colin Griffin, Andrew Steele, Norman Wainwright, Eileen Stansbery, 2000, Lunar and Planetary Science Conference XXXI
- [25] Kivelson, M.G., Khurana, K.K., Russell, C.T., Volwerk, M., Walker, R.J., Zimmer, C. 2000, Science, 289, 1340.
- [26] Figueredo, P.H., Greeley, R., Locating Potential Biosignatures on Europa from Surface Geology Observations, 2003, Astrobiology, 3, 851.

- [27] Figueredo, P.H., Greeley, R., 2004, Resurfacing history of Europa from pole-to-pole geologimapping, Icarus, 167, 287.
- [28] Dalton, J.B., Prieto-Ballesteros, O., Kargel, J.S., Jamieson, C.S., Jolivet, J., Quinn, R., Spectral comparison of heavily hydrated salts with disrupted terrains on Europa, 2005 Icarus, 177, 472.
- [29] Dalton, J.B., Mogul, R., Kagawa, H.K., Chan, S.L., Jamieson, C.S. Near-Infrared Detection of Potential Evidence for Microscopic Organisms on Europa, 2003, Astrobiology, 3, 505.
- [30] Biemann, K. 1979, J. Mol. Evol. 14, 65.70.
- [31] Biemann, K., Oro, J., Toulmin, P. III, Orgel, L.E., Nier, A.O., Anderson, D.M., Flory, D., Diaz, A., Rushneck, D.R., and Simmonds, P.G. 1977, *J. Geophys. Res.* 82, 4641.4658.
- [32] Mumma, M.J., Novak, R.E., DiSanti, M.A., and Bonev, B.P. A sensitive search for methane on Mars 2003, AAS, Division for Planetary Sciences meeting \#35, 1418 (2003)
- [33] Krasnopolsky, V.A.; Maillard, J.-P., Owen, T.C. Detection of methane in the martian atmosphere: evidence for life? Icarus, 172, 537 (2004).
- [34] Formisano, V., Atreya, S., Encrenaz, T., Ignatiev, N., and Giuranna, M. Detection of Methane in the Atmosphere of Mars, Science 306, 1758.1761 (2004).
- [35] McKay, C., Life on Mars, in *The Molecular Origins of Life*, ed. A. Brack, CUP (1998).
- [36] Bada, J.L., McDonald, G.D., Amino Acid Racemization on Mars: Implications for the Preservation of Biomolecules from an Extinct Martian Biota, *Icarus*, 114, 139 (1995).
- [37] Houtkooper, J.M., Schulze-Makuch, D., A possible biogenic origin for hydrogen peroxide on Mars: the Viking results reinterpreted http://arxiv.org/ftp/physics/papers/0610/0610093.pdf (2007).

Figure captions:

- Fig. 1. Hinds dual PEM polarimeter installed in the laboratory at NIST. The polarimeter views vertically down through an open port at the top of an integrating sphere. The sample is directly viewed at an open port below the sphere and for relfection experiments, illustrated to the right, a source of illumination enters a third port onto the interior of the sphere without direct illumination of the sample. The net result is illumination of the sample by diffuse unpolarized light. Two examples of cyanobacteria samples in suspension are shown lower right.
- Fig. 2. Polarization analysis for cyanobacteria WH8101. Fig. 2A shows the shallow reflection circular polarization spectrum with ±1-σ error bars. Overlaid for comparison, dotted line, is the data from the transmission experiment. Above, are arbitrarily scaled representations of the absorbance (solid line) and degree of linear polarization (dashed line) to illustrate spectral relationships between absorption features and circular polarization features. In general the degree of linear polarization is much smoother than the circular polarization. See Sparks et al (2009). Fig. 2B shows the concentrated microbial samples, which have lower S/N given they are darker. Overlaid are the shallow reflection (solid line) and transmission (dotted line) polarization spectra. Above, the absorbance (solid line) and linear polarization (dashed line) spectra with arbitrary scaling.
- Fig. 3. The average circular polarization reflection spectrum of a miscellaneous sample of macroscopic vegetation showing significant levels of polarization throughout, but especially in the chlorophyll absorption band between 650 and 700 nm. An example arbitrarily scaled reflectance spectrum is shown below.
- Fig. 4. Circular polarization reflection spectra of control minerals (filled circles with error bars) and arbitrarily scaled reflection spectra (solid line below). Fig. 4A shows the spectrum of

a sulphur rock with a prominent absorption edge in the blue. Fig 4B shows the Mars regolith stimulant JSC-1 which has a smooth spectrum. While non-zero polarization levels are measured, they do not show any correlation with features in the intensity spectra.

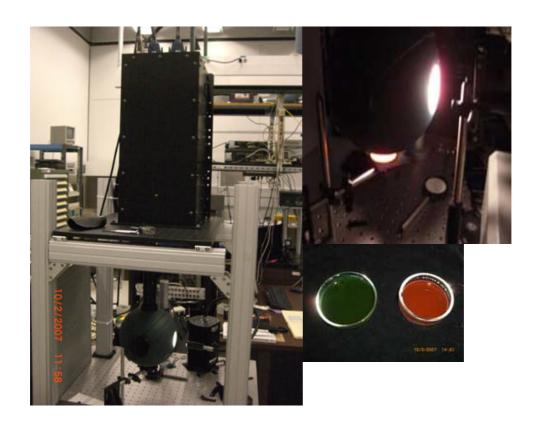


Fig. 1

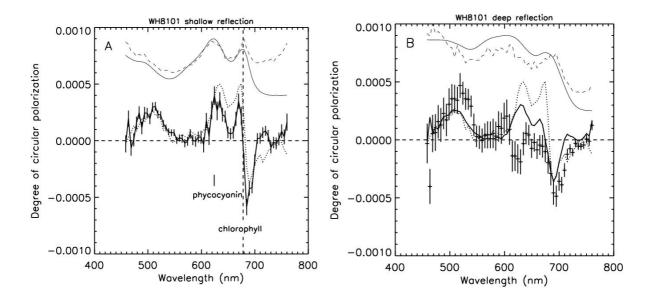


Fig. 2

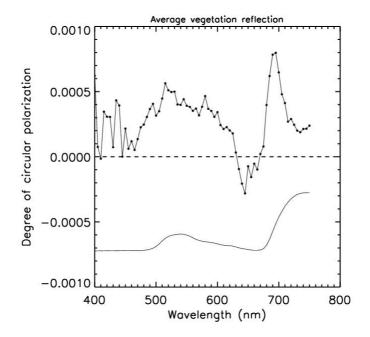
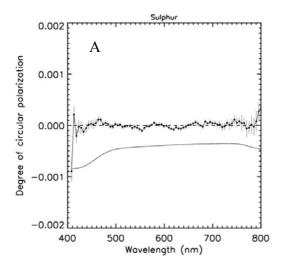


Fig. 3



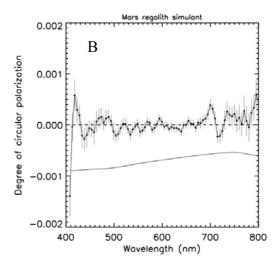


Fig. 4